# Machine Learning Prediction of Pebble History and Nuclide Concentration in PBRs

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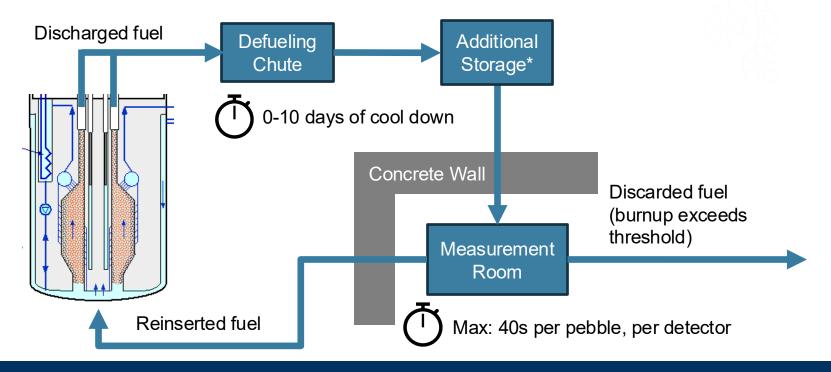


#### Pebble Assessment

- Objective: Simulate a realistic detected gamma spectrum from a gFHR fuel pebble and use ML to make predictions with it
- Challenges and Considerations:
  - About 2,300 pebbles leave the core a day, giving us about 40 seconds max to measure each pebble
  - Even with decay times of up to several days, pebbles emit on the order of 10<sup>13</sup> gammas per second
  - For a detector, this could lead to serious dead time without prohibitively expensive shielding
- Can ML provide an edge over regression approaches?



## Hypothetical Burnup Measurement System



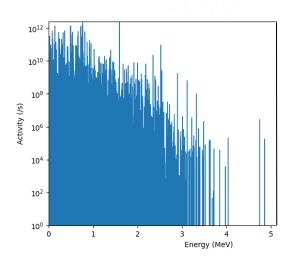


# Detected Spectrum Generation



### **Emitted Gamma Source Generation**

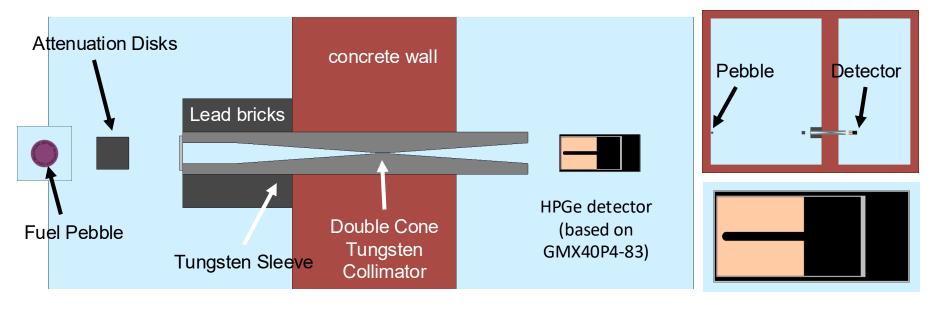
- Discharge compositions used to generate spectra
  - Unique, pebble-level data from HxF
  - Kairos Power gFHR at equilibrium
- Different decay times between 0 and 10 days
  - 120 second variance applied
- Spectra produced in Serpent
  - Decay step (neutron mode)
  - Gamma source (photon mode)





#### Detector Model (HPGe)

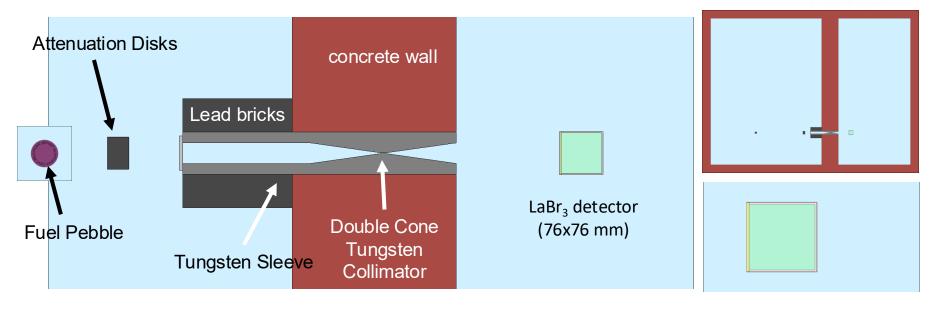
Total pebble distance	250 cm	
Collimator Length	54 cm	
Detector Dimensions	6.06 x 6.69 cm	
Photopeak Efficiency @ 1MeV	2.19 x 10 <sup>-8</sup>	





#### Detector Model (LaBr3)

Total pebble distance	170 cm	
Collimator Length	27 cm	
Detector Dimensions	7.6 x 7.6 cm	
Photopeak Efficiency @ 1MeV	1.87 x 10 <sup>-7</sup>	





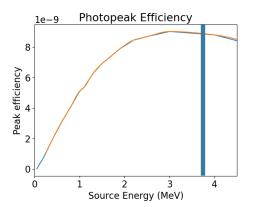
# Detector Response Simulation: Tally

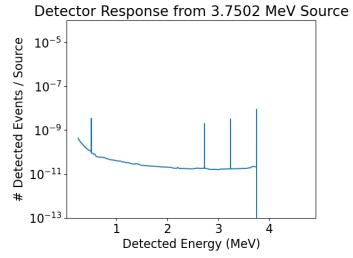
- Simulated at 16 different source energies
- Monoenergic source sampled from pebble
- Pulse height tally in detector crystal
- Energy bins in the tally are divided to capture different reactions:
  - Coarse bins (20-50 keV wide) for continuum
  - Very fine bins for discrete energies (photopeak, pair production, single and double escape peaks)



# **Detector Response Function**

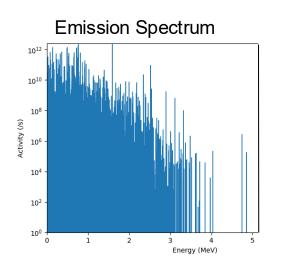
- 12000 channels from 0.25 MeV to 4.7 MeV
- Discrete peaks efficiencies are interpolated
- Continuum is fit with a RFR model

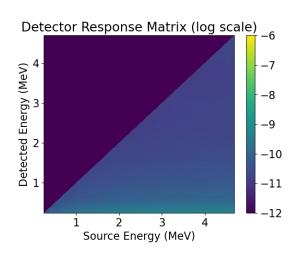


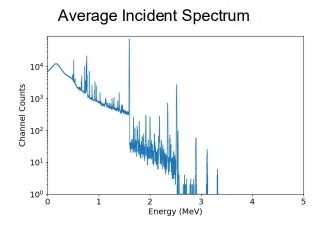




# Detector Response Matrix









# Dead time and pile-up

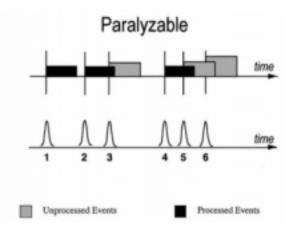
40s measurement time with 0.4s standard deviation With average count rate *R* and measurement time *t*:

- Sample total counts N~Poisson(Rt)
- Sample N energies E from average spectrum
  - Gaussian broadened
- Sample N waiting times t~Exp(1/R)

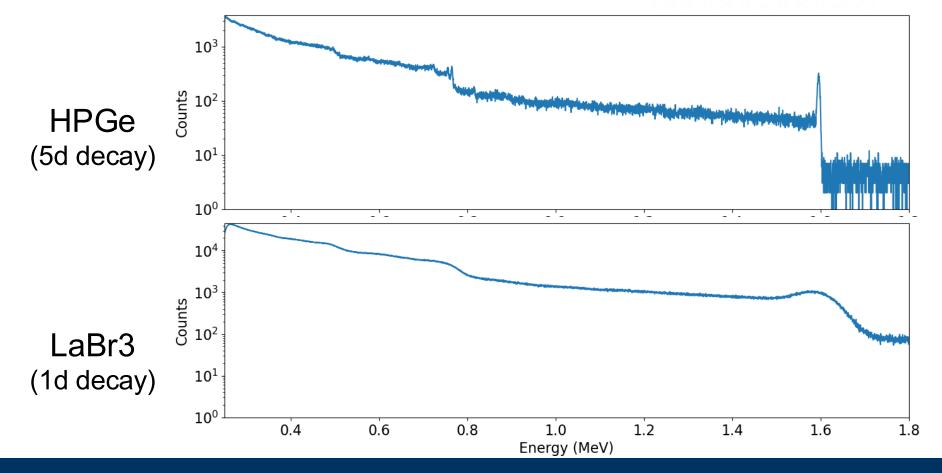
#### **Events arriving:**

- during peaking time are summed
- during dead time are rejected
- after are recorded normally

HPGe Dead Time	6 µs	
LaBr3 Dead Time	0.016 µs	





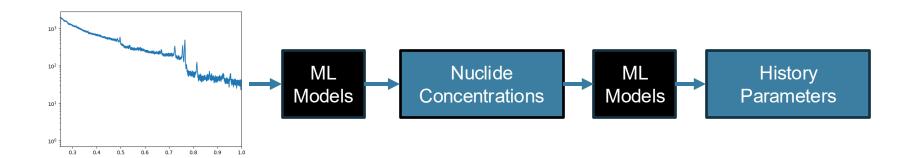




## **Pebble Prediction**



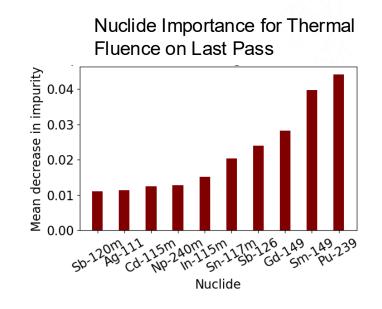
# Pebble Assessment: Two Layers





# Nuclide Feature Importance (Ideal)

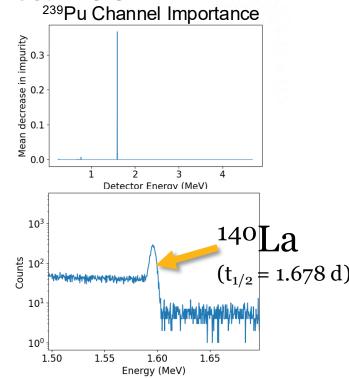
- A Random Forest Regressor (RFR) model is trained for each history parameter using actual concentrations from Serpent
- Informs which nuclides we need to predict
- Highlights physical trends and correlations





Spectrum Feature Importance

- For each nuclide, a RFR is fit using the actual spectral data
- The most important channels are determined
- Then, a NN model is trained and its hyperparameters optimized using only useful channels
- Rarely determined by that nuclide's own decay gammas





# Final Model Training

- Neural Networks are optimized for each history parameter using the predicted important nuclides
- Number of hidden layers and layer sizes are coarsely optimized in large grid search

#### Random Forest Regressor Parameters

Max Depth	14
# Estimators	1000
Min. Feature Split	1
Bootstrapping	True

#### **Neural Network Parameters**

# Epochs	300
Batch Size	256
Activation	ReLU
Learning Rate	0.001
Network Sizes tried	(8, 2), (64, 8), (128, 16), (36)



# **History Metrics Performance**

R<sup>2</sup> accuracy of model (HPGe – 40s measurement – 5d decay)

	LR on full Spectrum	NN on pred. nuclides	RFR on actual nuclides
%FIMA (burnup)	0.8021	0.9740	1.0000
Residence Time	0.8003	0.9671	0.9999
Passes	0.7974	0.9691	0.9999
Avg radial path on last pass	-1.9720	0.6643	0.9554
Thermal fluence on last pass	-3.9159	0.5576	0.8468

Baseline

Detected

Idealized



# History Metrics Performance

R<sup>2</sup> accuracy of model (LaBr3 – 40s measurement – 1d decay)

	LR on full Spectrum	NN on pred. nuclides	RFR on actual nuclides
%FIMA (burnup)	0.7770	0.9856	1.0000
Residence Time	0.5851	0.9888	0.9999
Passes	0.5825	0.9848	0.9999
Avg radial path on last pass	-1.8001	0.8969	0.9554
Thermal fluence on last pass	-6.2175	0.8369	0.8468

Baseline

Detected

Idealized



## Conclusions & Final Work

- Conclusions
  - Machine learning offers greater accuracy than linear regression
  - Timing resolution is more important than energy resolution
  - HPGe detectors may be prohibitively expensive to run
- Future Work
  - Explore other detector choices & further optimize setup
  - Run different measurement times (multiples of 40s)
  - More extensive ML model optimization



# Thank you!

#### Acknowledgements

- Professor Kai Vetter for insight on detector setup and modeling
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